

the

H A L O

Monitor



DEELS , 12-13 May 2014

The vertical beam Halo monitor

1) the Halo itself : what is it ?
 what creates it ?
 how does its vary with different conditions ?

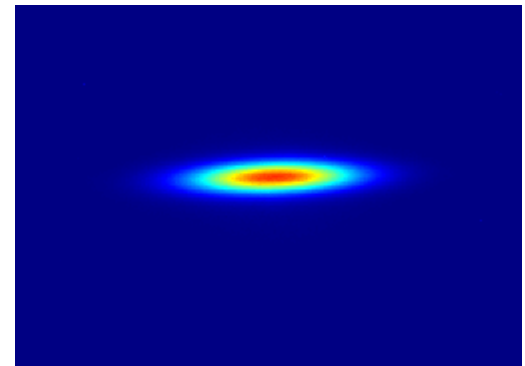
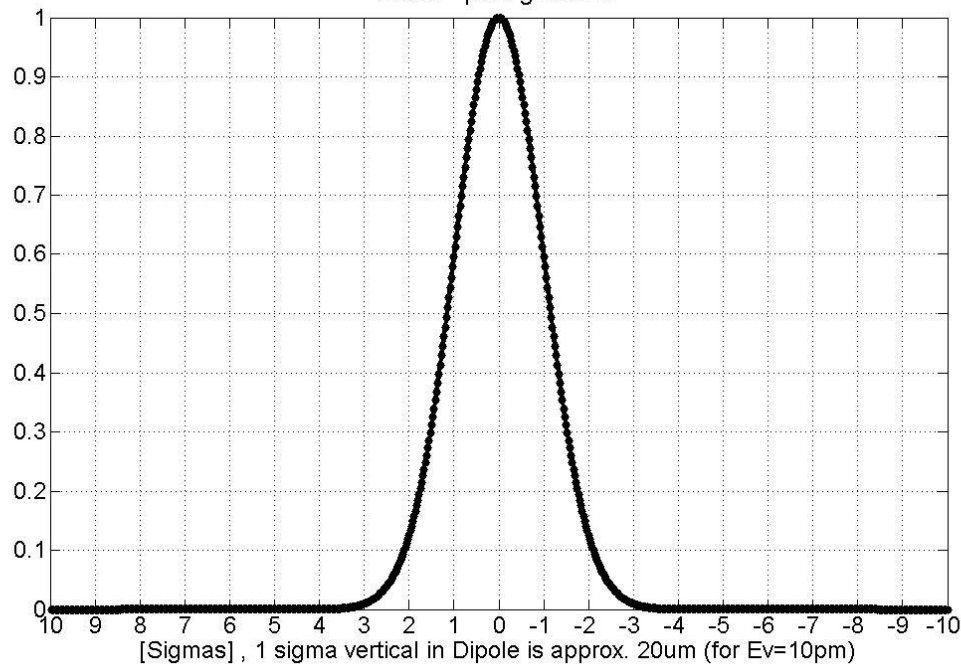
2) Principle of technique :

- using X-rays from a dipole
- on a beamport with adapted absorbers
- with a light-blocker to shadow against the intense beam-core
- and a sensitive detector further down-stream

3) Results, Calibration, Initial Problems & Solutions, Conclusions

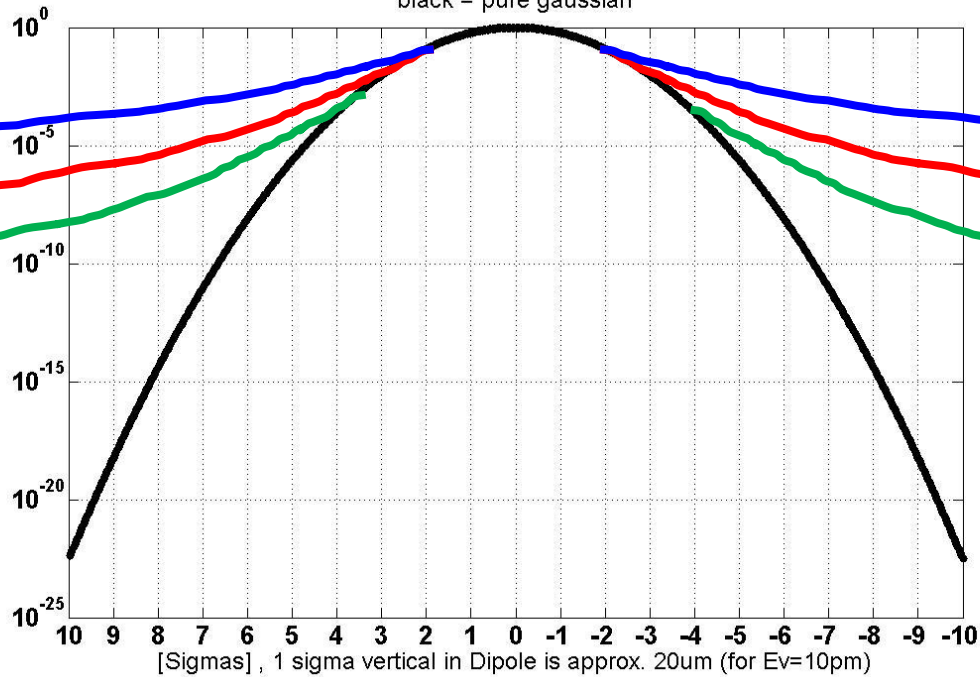
black = pure gaussian

Linear Scale

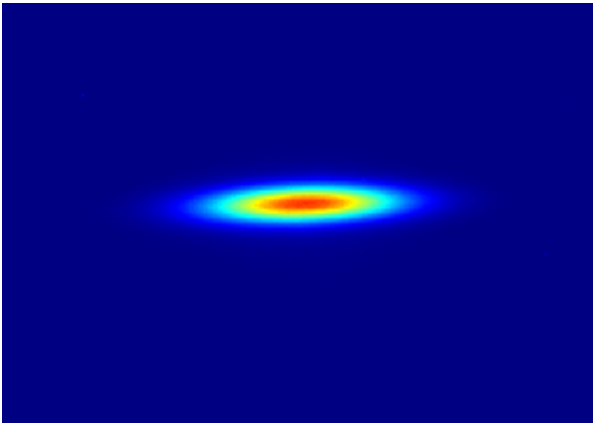
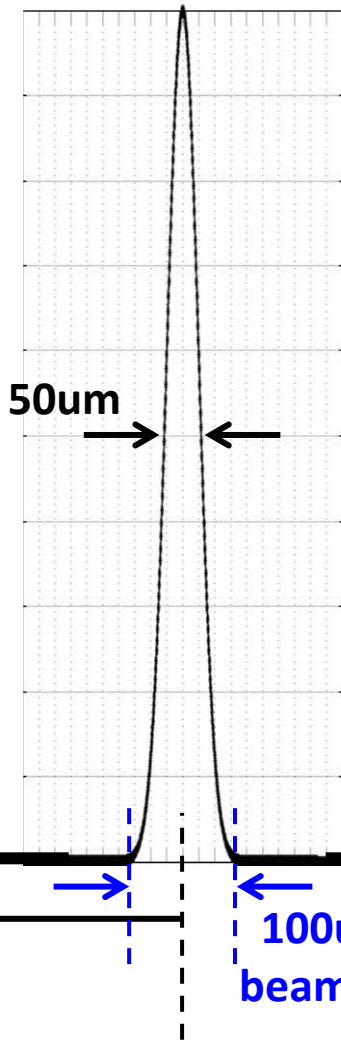


black = pure gaussian

Log. Scale



??
??
??



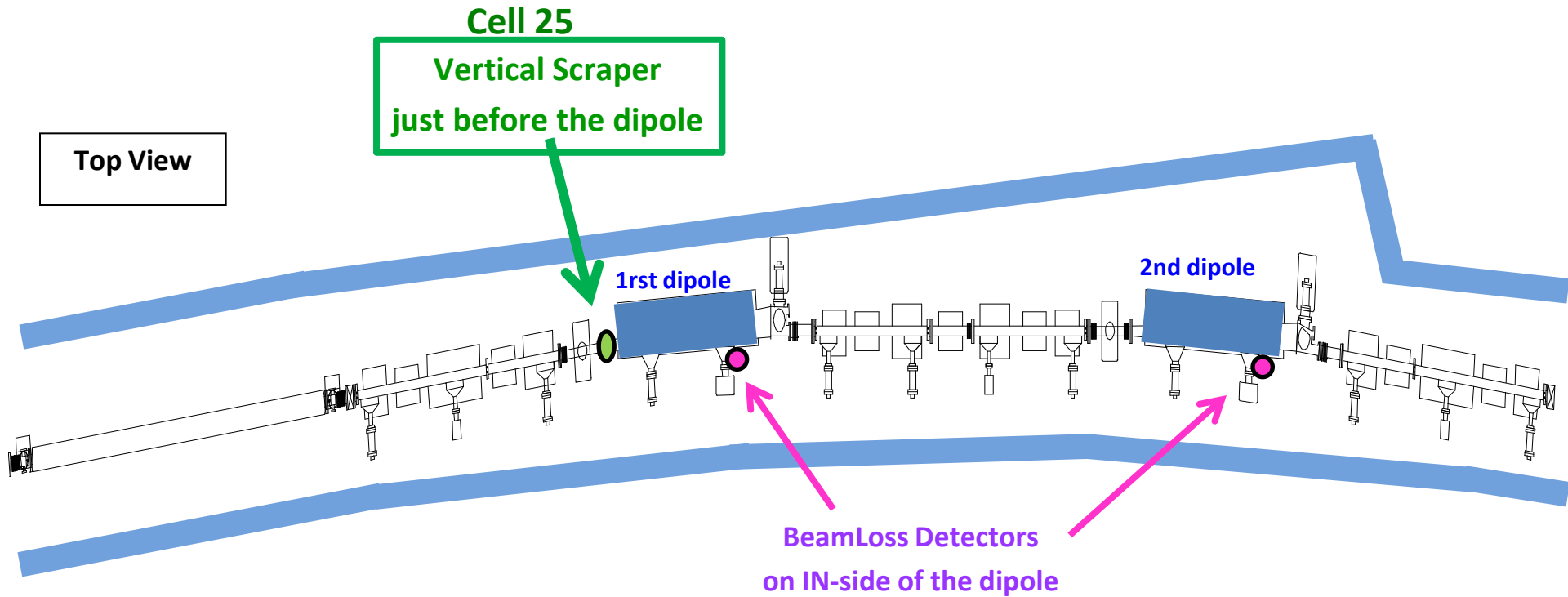
←
upto
10mm

←
100um =
beam footprint

i.e. 100 x footprint

there are still electrons

Scraper Manipulations



In total in the Ring :

1 horizontal scraper (C4)

3 vertical scrapers (C5, C22, C25)

Scraper material & thickness very different :

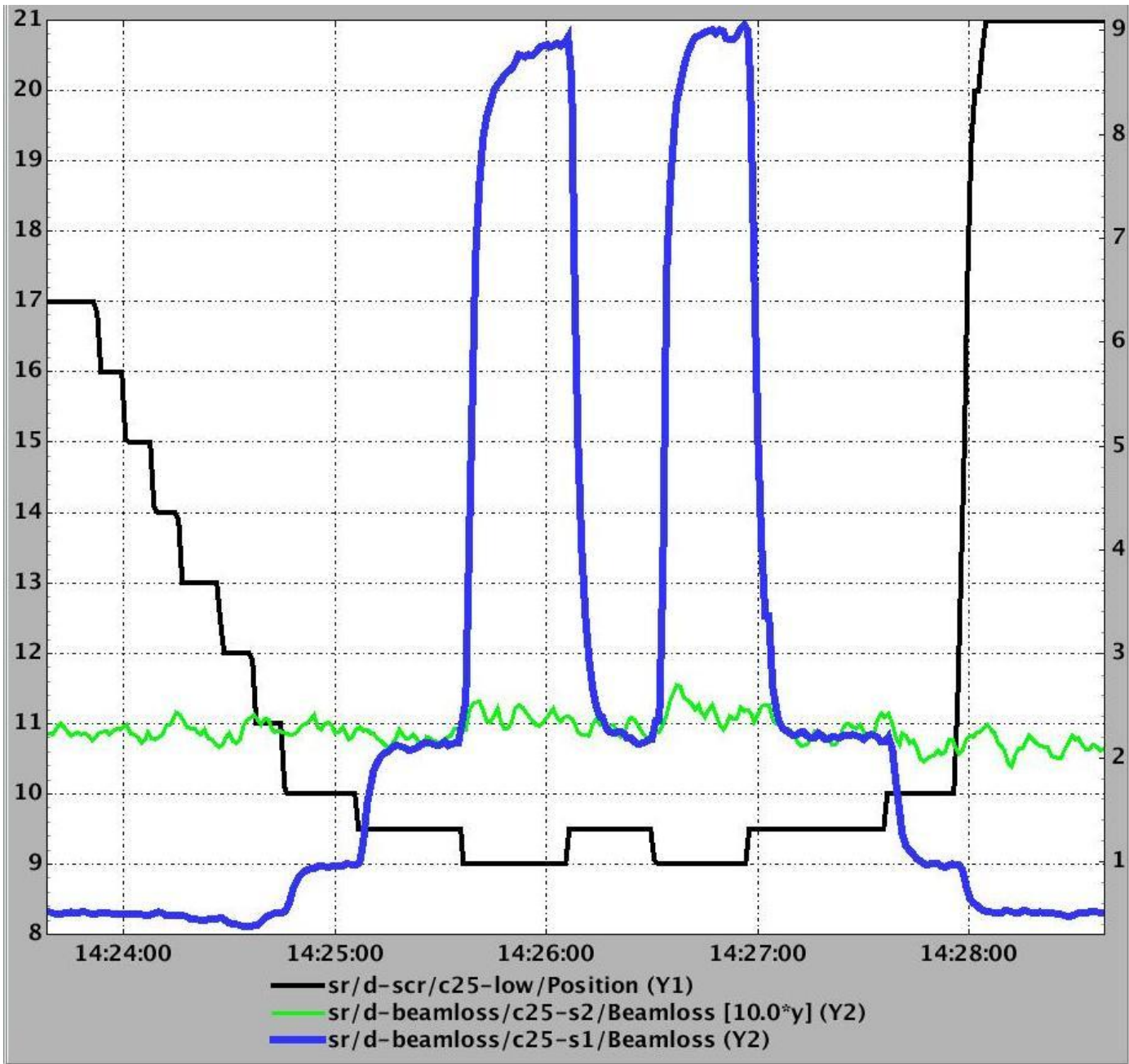
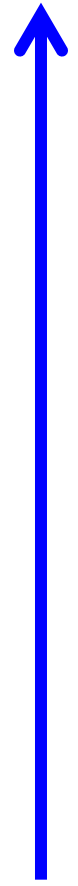
Copper : 20mm, 30mm, 90mm

Tantalum : 90mm

Black :
Scraper
position
[mm]



Blue :
BLD-signal
[a.u.]



The vertical beam **Halo** monitor

what is the Halo ? **Electron density, at “large” distance from the central core**

what creates it ? **Scattering between :**
- the electrons themselves (Touchek)
- electrons & residual gas particles

how does its vary with different beam conditions ?

- stronger beam density → stronger Halo
- poor vacuum → stronger Halo

The non-destructive vertical beam Halo monitor

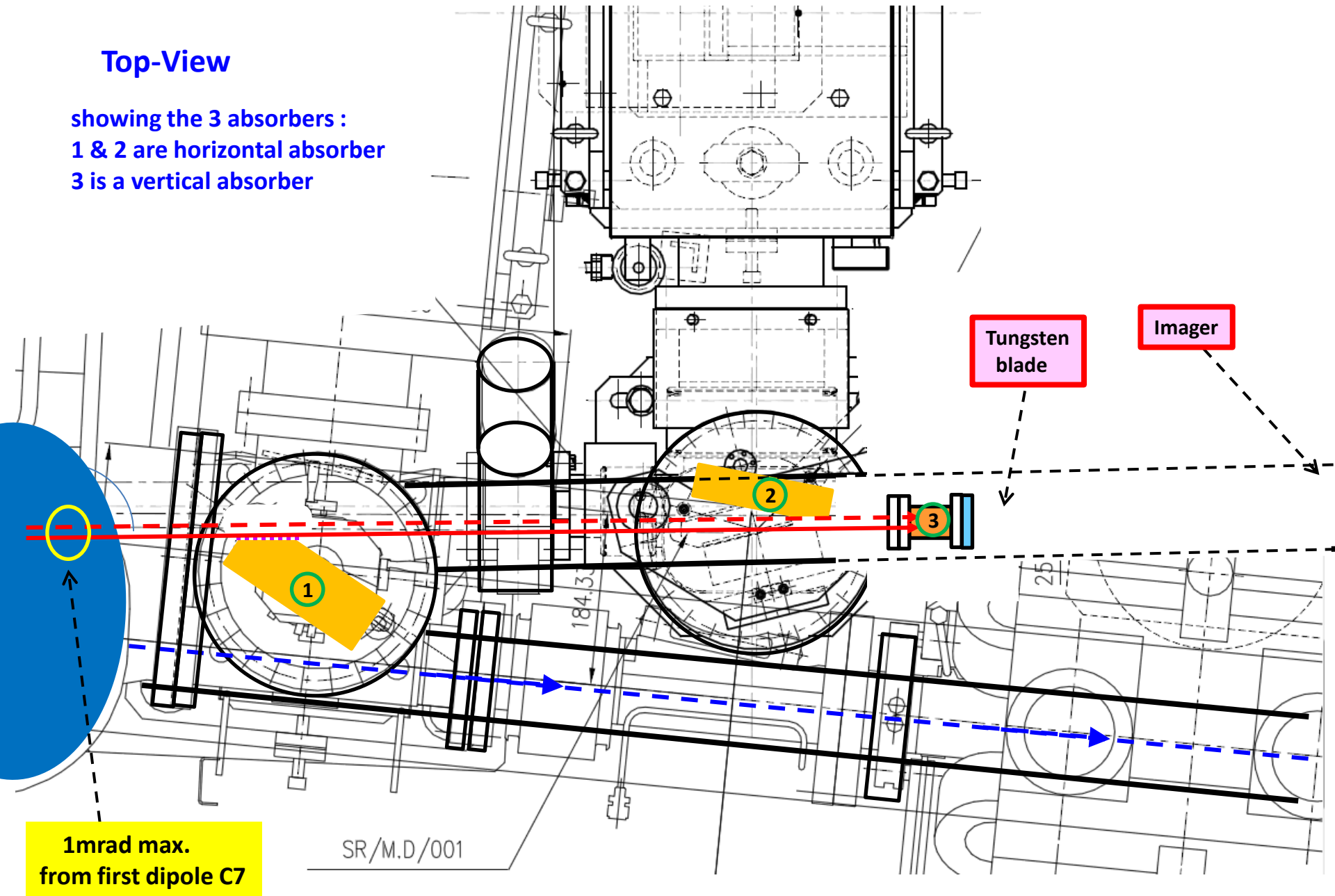
Principle of technique :

- using X-rays from a dipole
- on a beamport with an adapted absorber
- with a light-blocker to shadow against the intense beam-core
- and a sensitive detector further down-stream

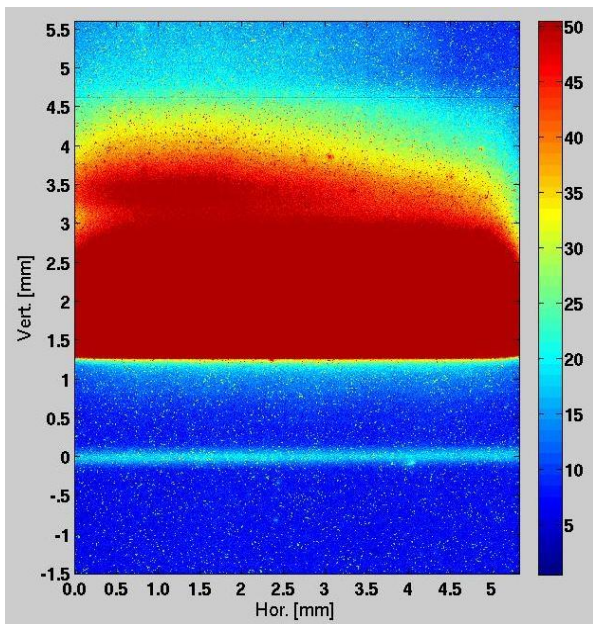
Goal : the permanent monitoring of the Halo intensity,
in a region of 2 – 10mm above the beam-core,
and with a calibrated value of this intensity

Top-View

showing the 3 absorbers :
1 & 2 are horizontal absorber
3 is a vertical absorber

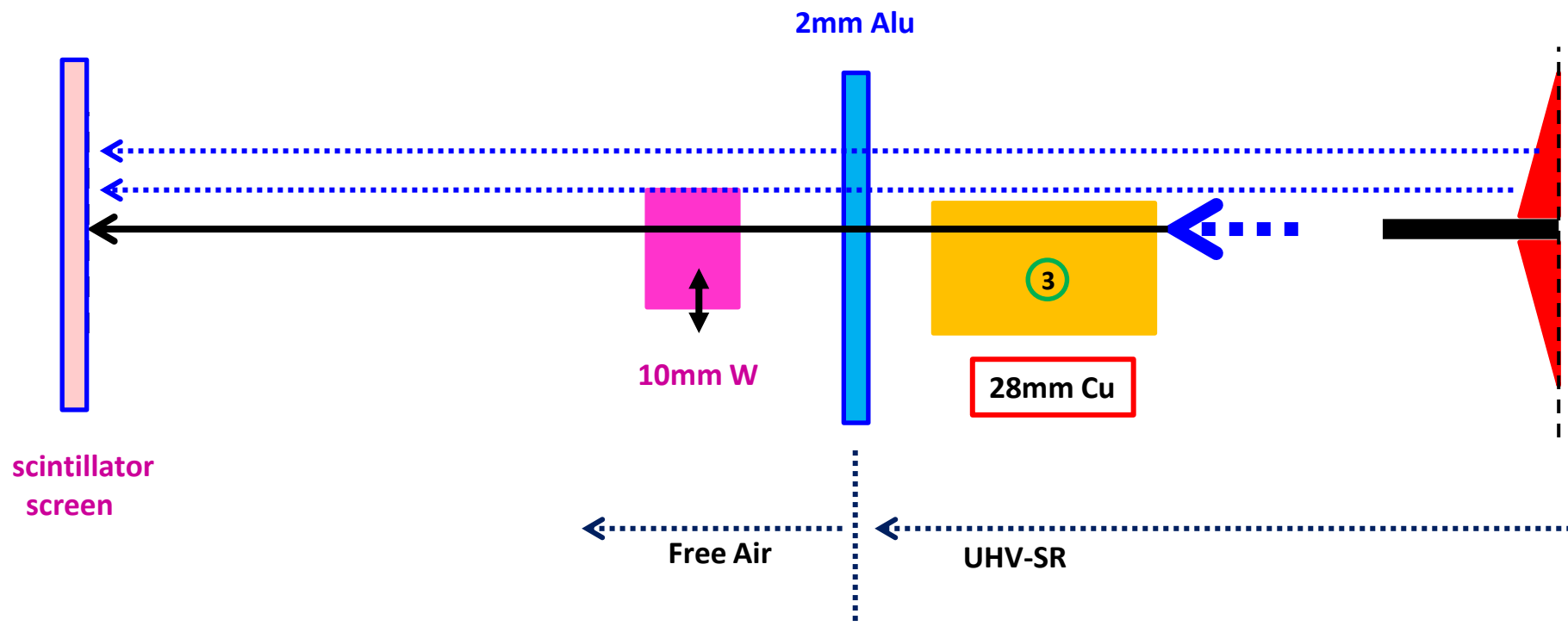


Side-View



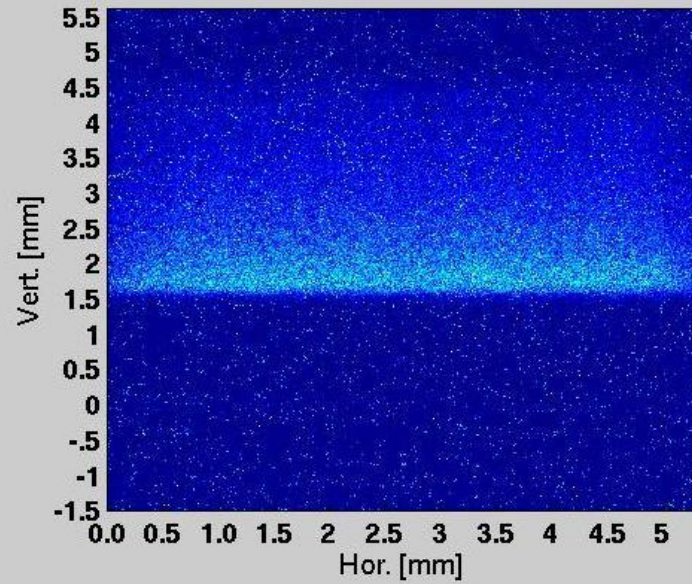
←..... is this the Halo ? or is it an artefact (scattering, reflection) ??

←..... from the beam-core (very strongly attenuated)

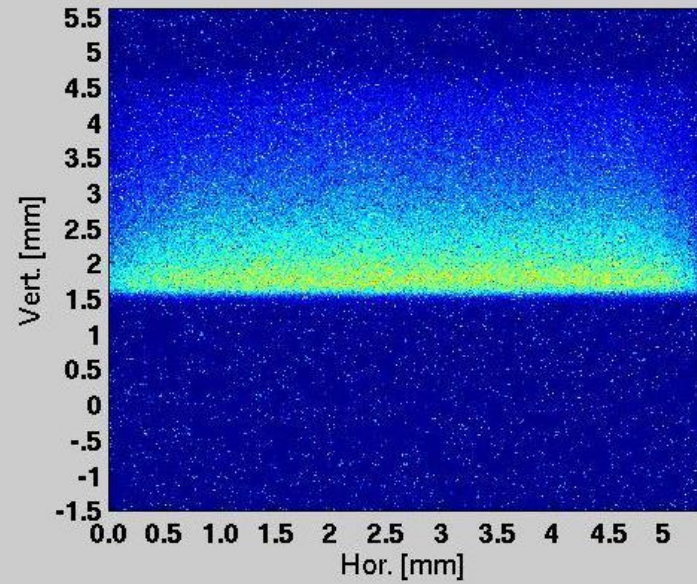


Differential Images while scraping

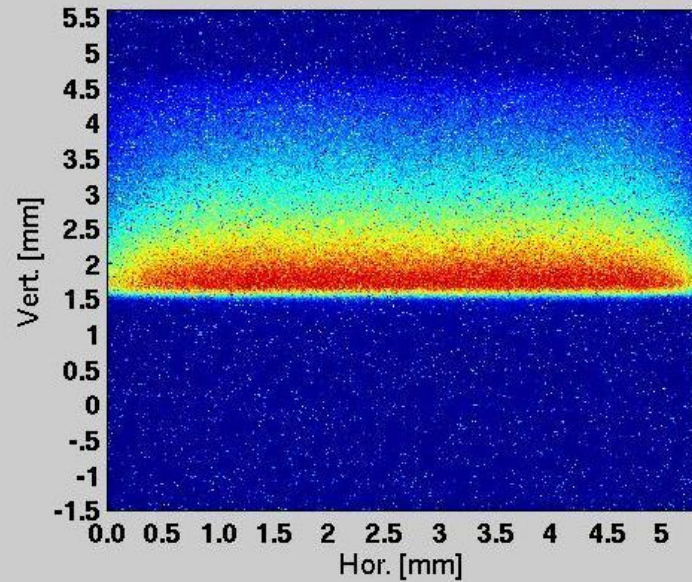
scraper at 6 mm



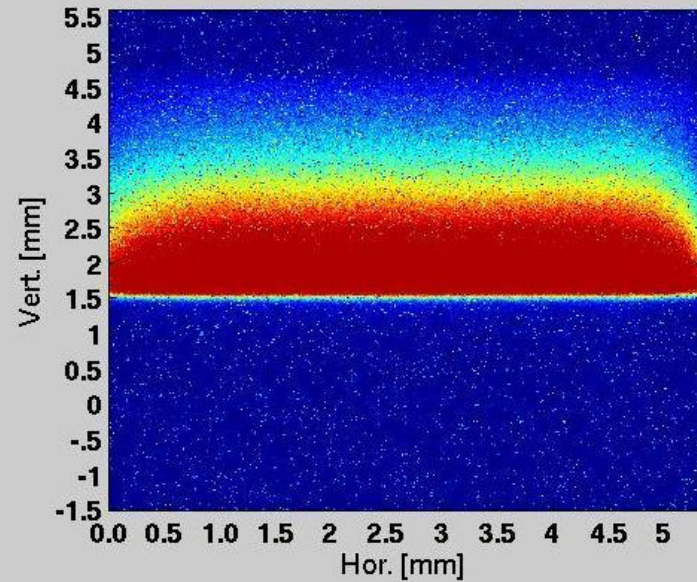
scraper at 5 mm



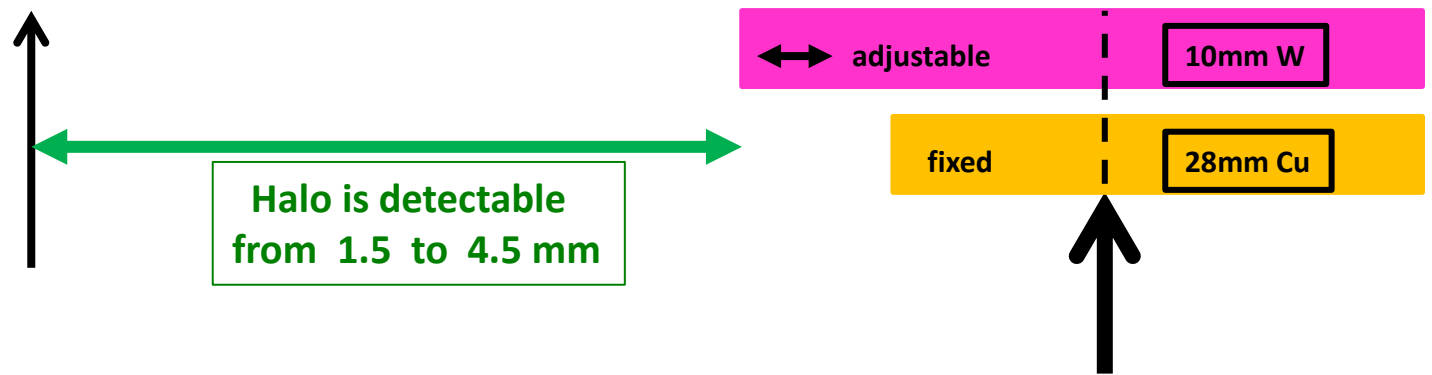
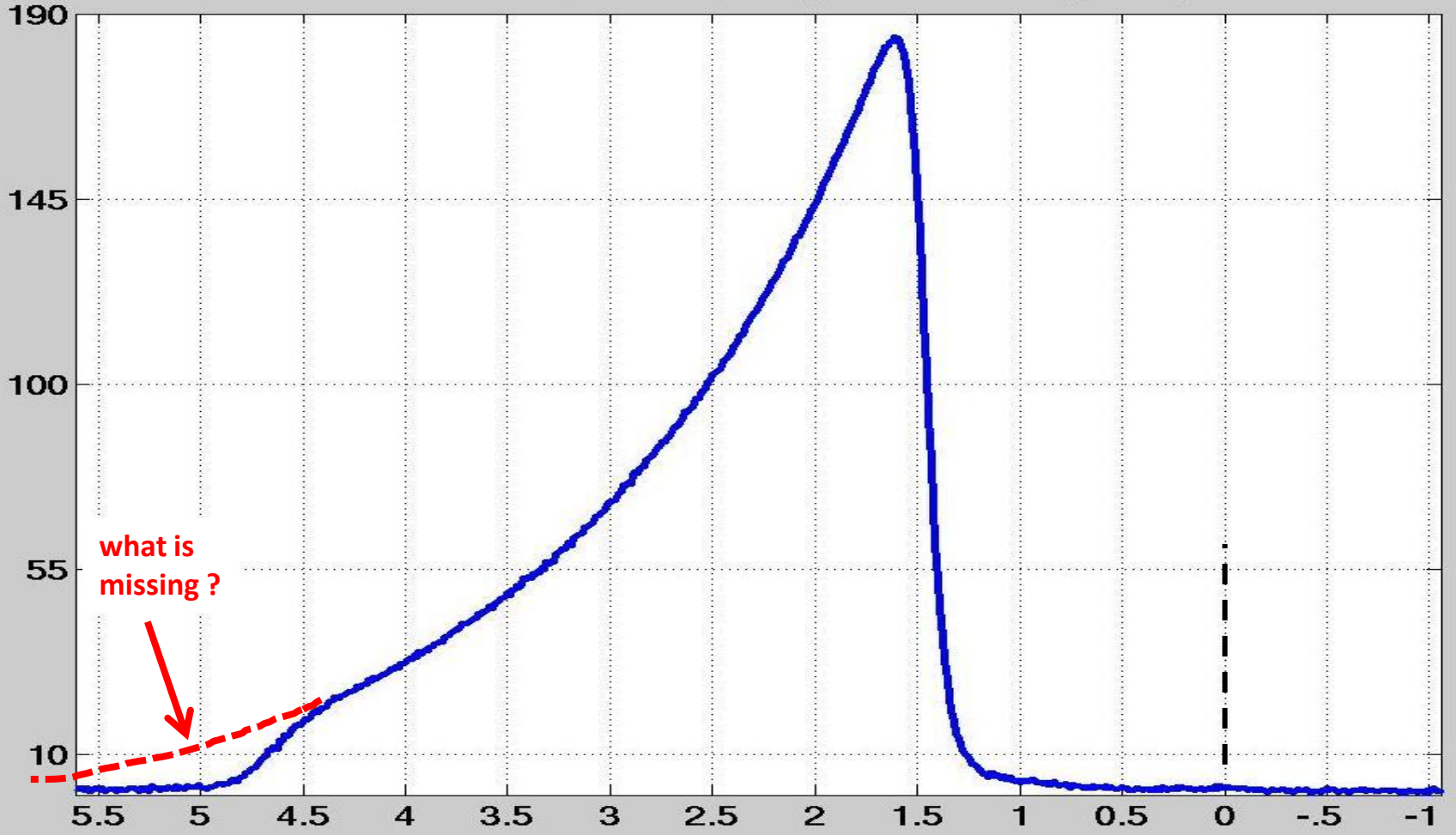
scraper at 4 mm



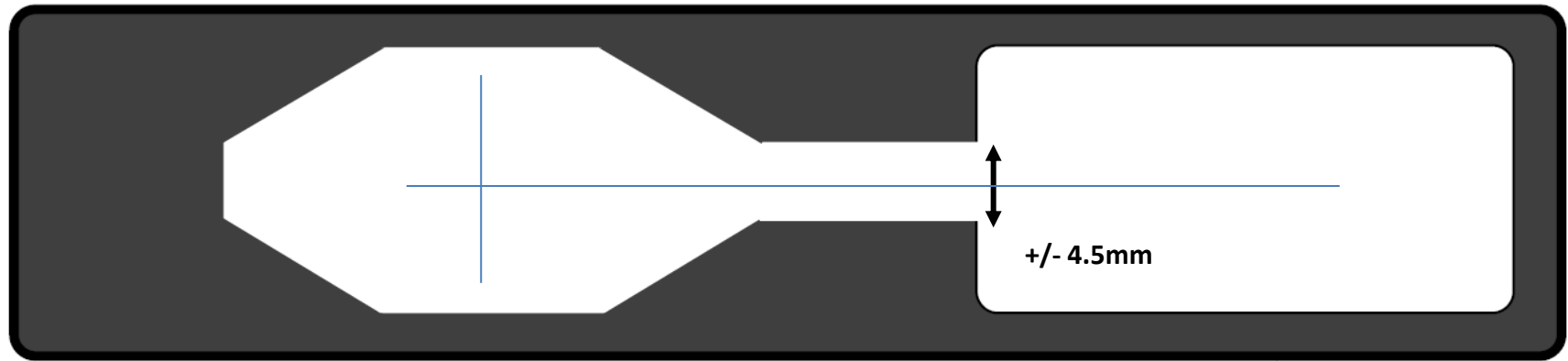
scraper at 3 mm



differential profile (scraper In & fully Out)



Dipole chamber cross-section



↑
Electron
Beam
Chamber

↑
RF shielding
Standard gap = 9 mm

↑
Photon
Beam
Chamber

Results with various Beam manipulations :

- 1) Closing the In-Vac undulators : to verify if (and how much) they “scrape” the beam**
- 2) Detecting the fluctuations of the UHV quality (detecting ultra-small gas-outbursts)**
- 3) Measuring the Halo strength while varying the Touchek scattering conditions**

**Can the Halo monitor now measure the effect
of closing the Gap of the In-Vacuum undulators ??**

**We assessed this precisely by closing the gap of the In-Vacs
of ID-11 and of ID-15 from 8 to 6mm (in 0.2mm step)**



see results in next slides

ID-15 In-Vacuum

Differential images while closing & opening the Gap from 8 to 6 to 8 mm

8 mm



7.8 mm



7.6 mm



7.4 mm



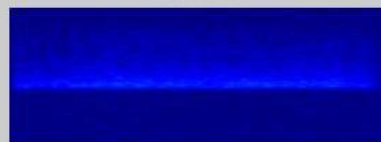
7.2 mm



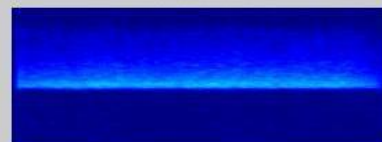
7 mm



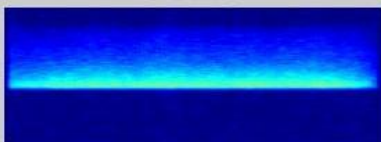
6.8 mm



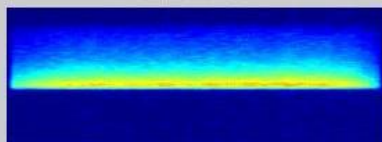
6.6 mm



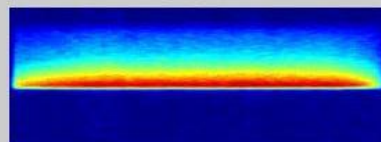
6.4 mm



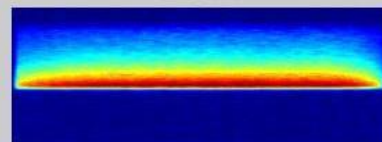
6.2 mm



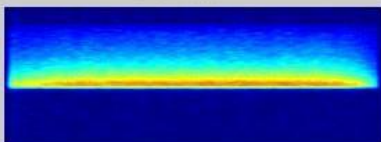
6 mm



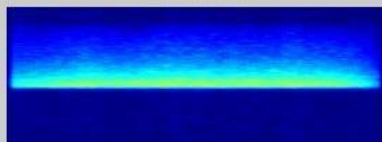
6 mm



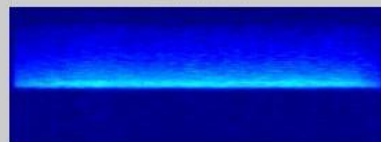
6.2 mm



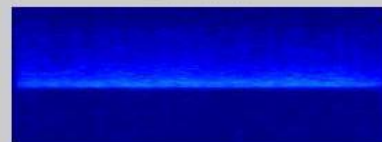
6.4 mm



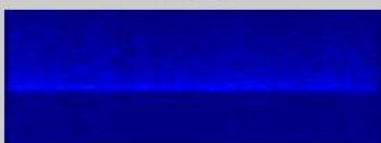
6.6 mm



6.8 mm



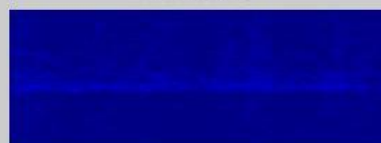
7 mm



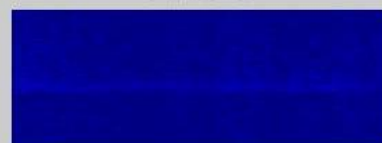
7.2 mm



7.4 mm

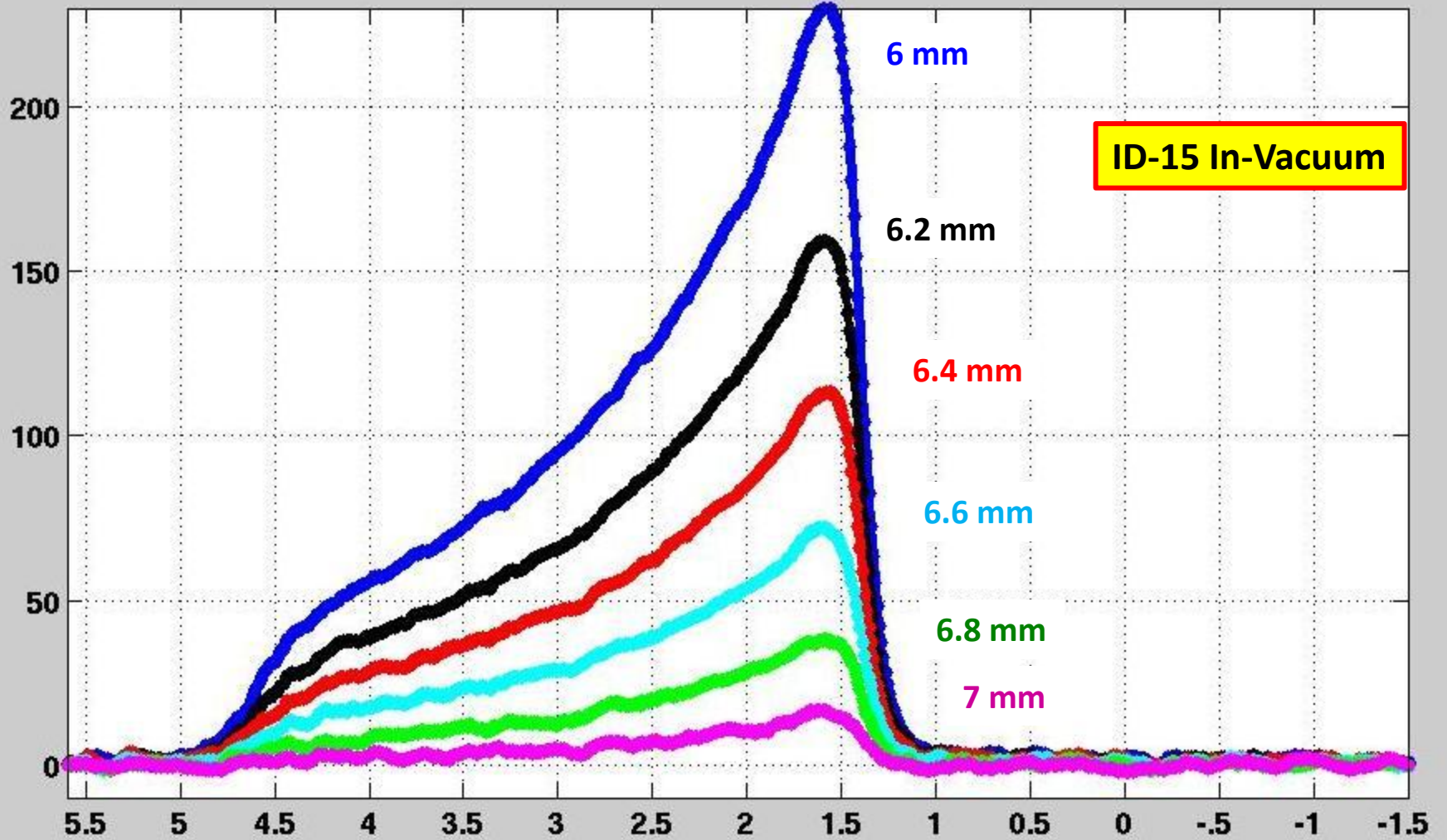


7.6 mm



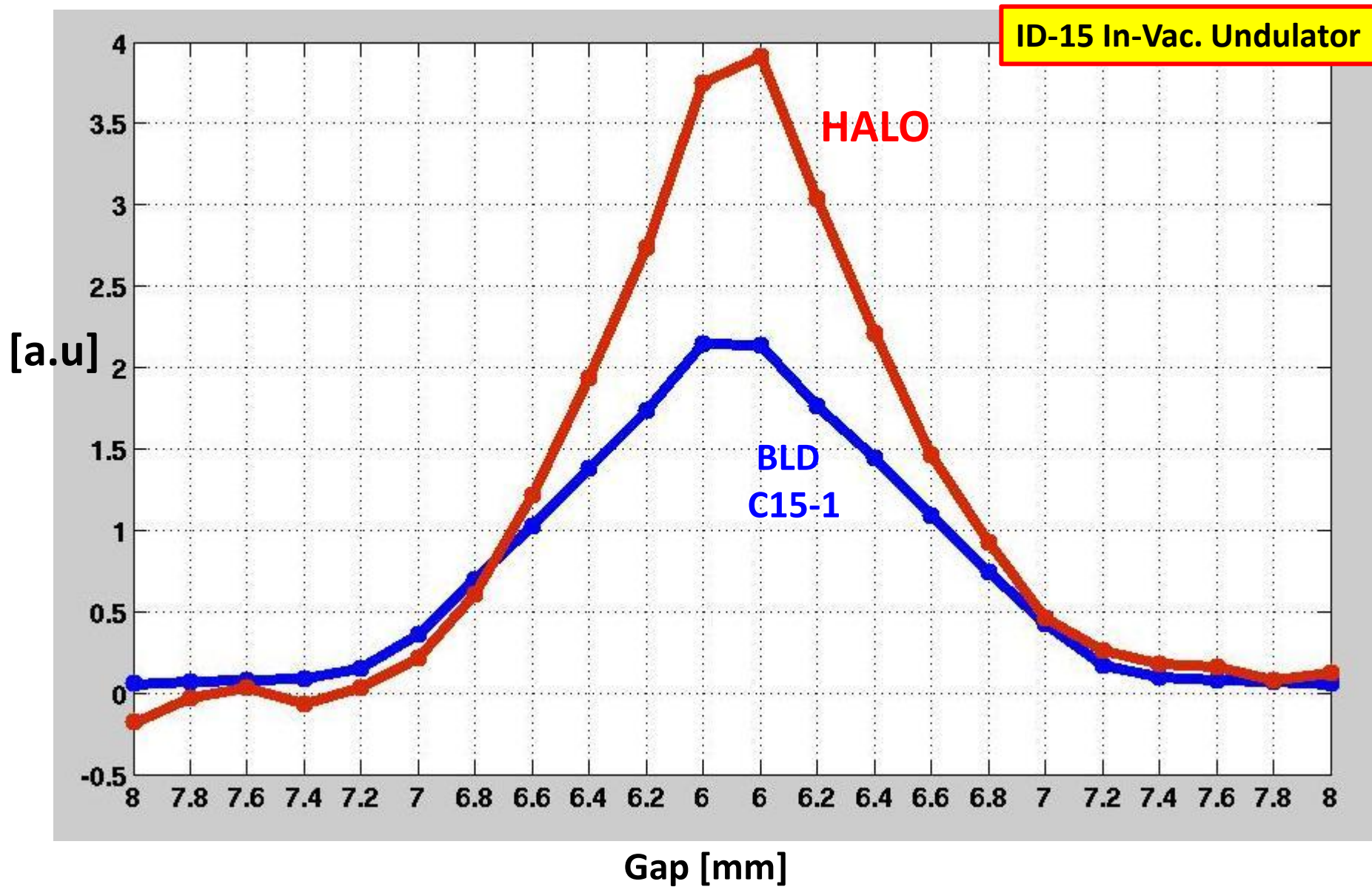
Gap from 7 to 6 mm

ID-15 In-Vacuum



← Distance from beam centre [mm]

how does the **Halo monitor** compare with the (local) **BeamLoss Detector** :

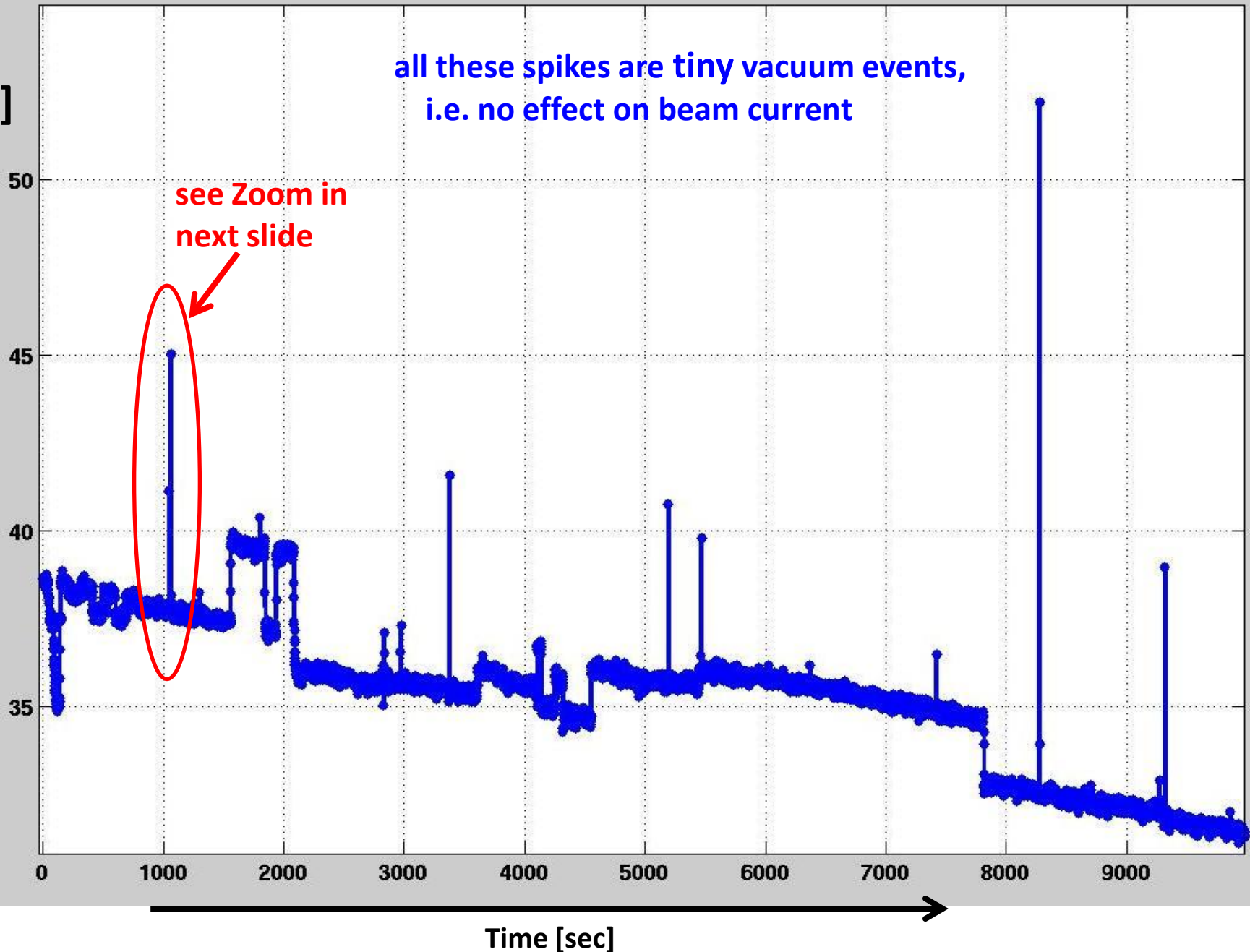


Halo signal over 3 hours during USM

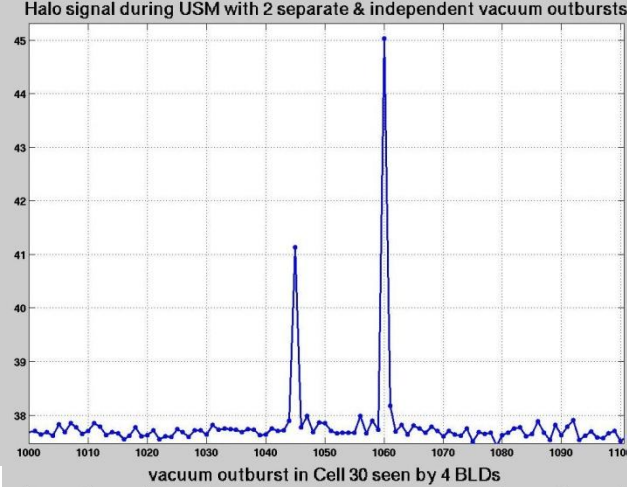
[a.u]

all these spikes are tiny vacuum events,
i.e. no effect on beam current

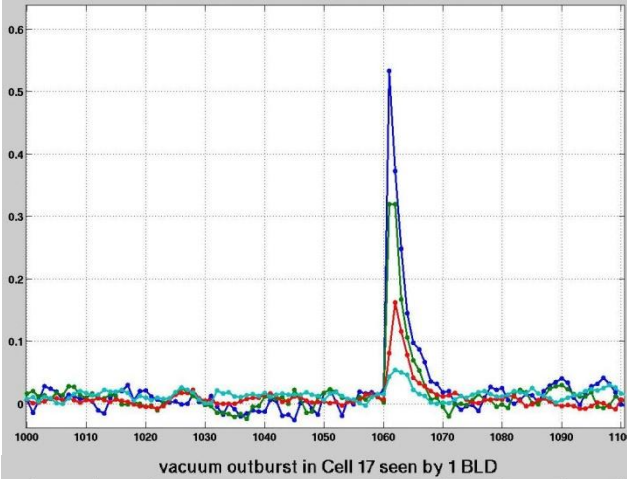
see Zoom in
next slide



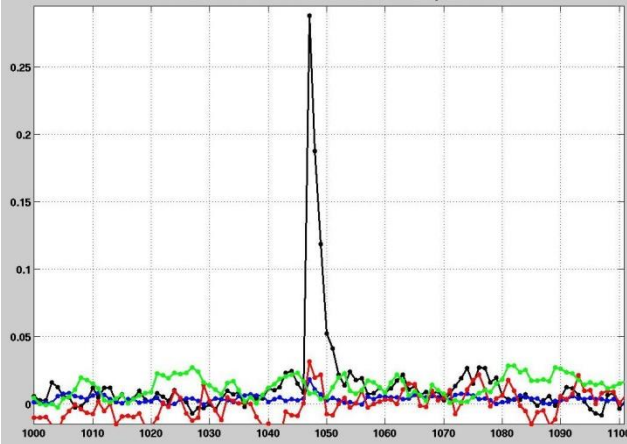
Halo



**BLDs
C30-31**



**BLD
C17**



**the Halo is very sensitive
to vacuum quality :**

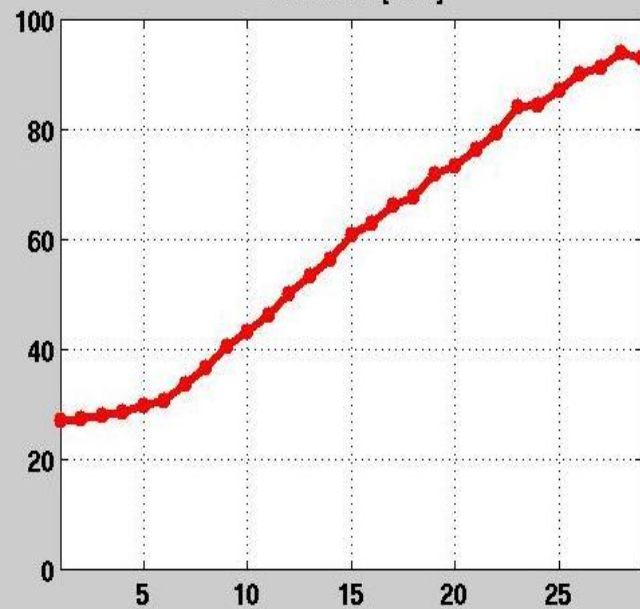
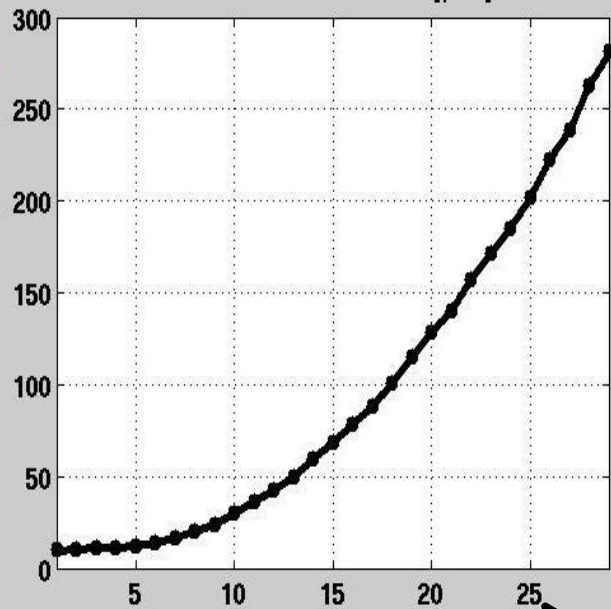
**in terms of sensitivity
and signal/noise ratio**

**the Halo is comparable
to a BLD that is close
to that Vacuum event**

**so the Halo is the
best “Watch-Dog” of
general SR vacuum quality**

all data @1Hz, 1000sec

Lifetime [hrs]

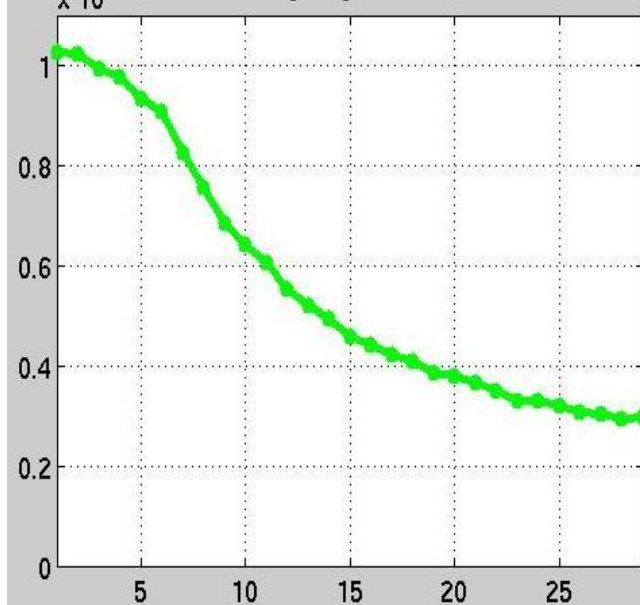
Vertical Emittance [μm]

Comparative results when :

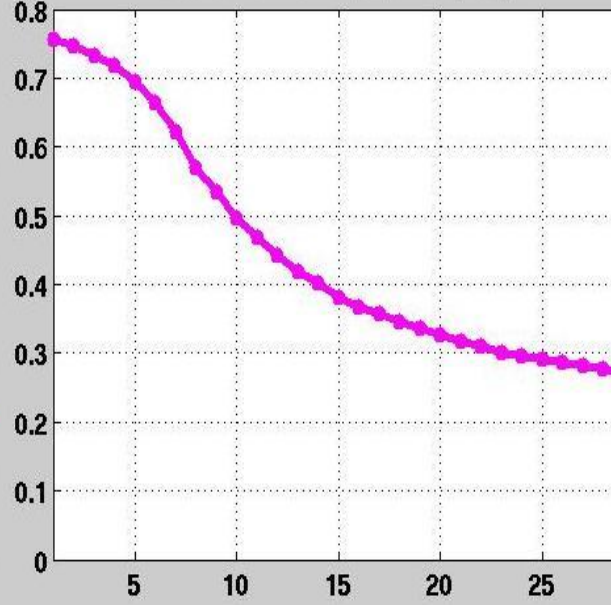
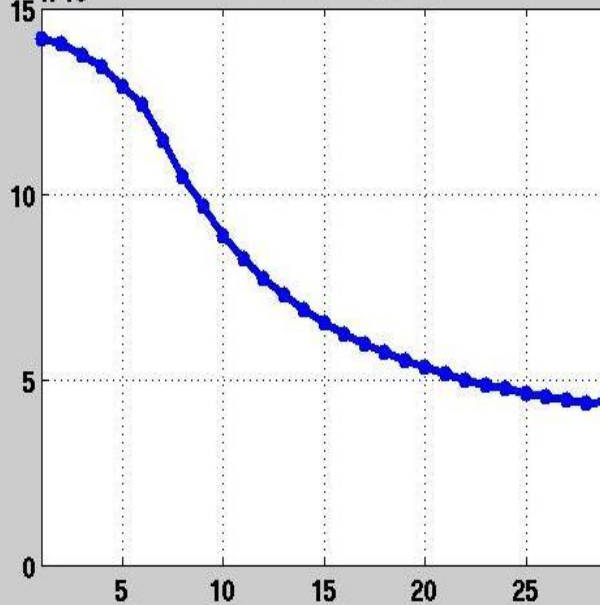
increasing the vertical beam size

thus reducing the Toucek effect
and going towards
(mainly) UHV effect

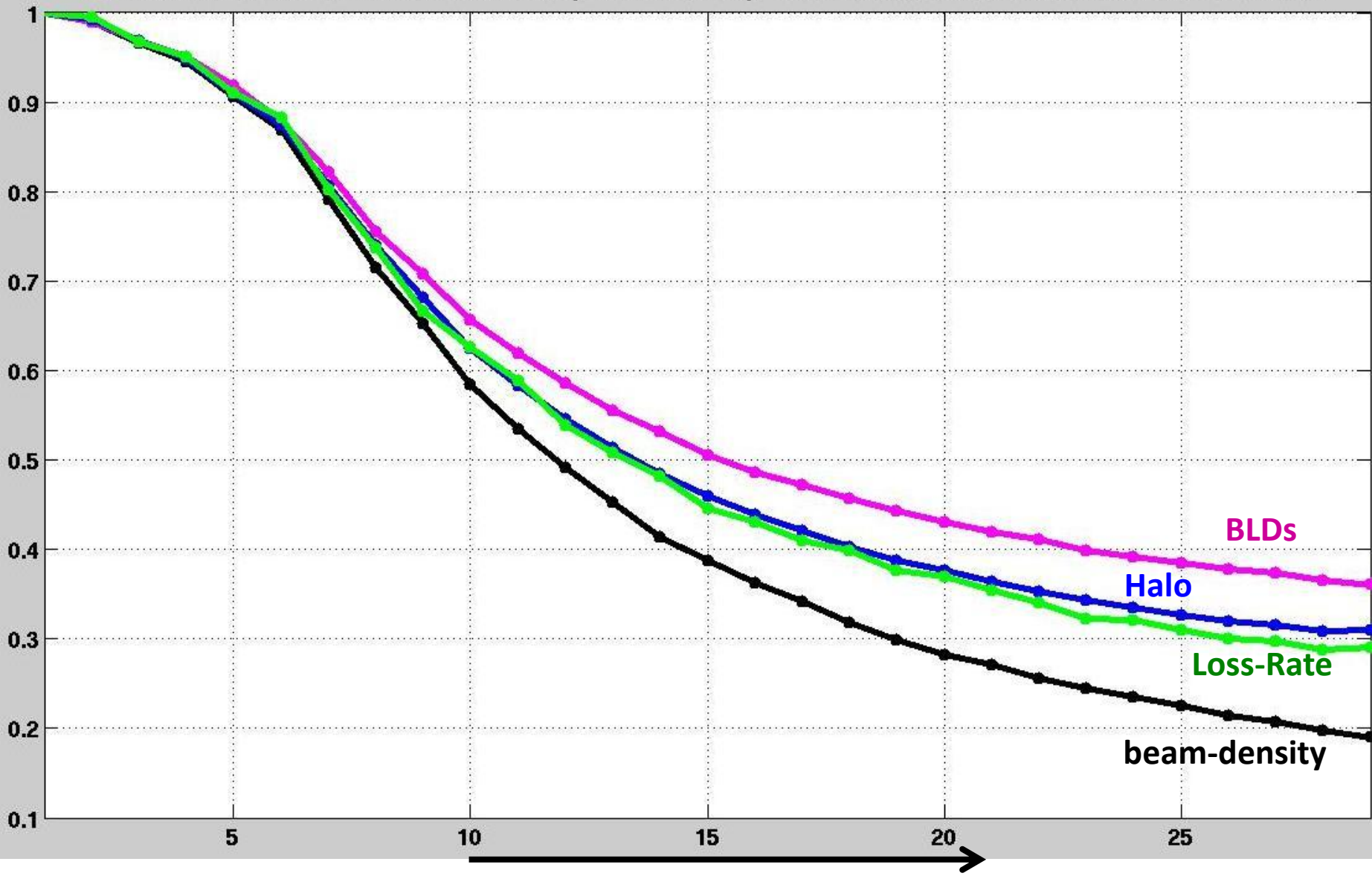
increasing vertical beam-size,
away from Toucek dominated scattering

 $\times 10^{-5}$ Loss Rate [a.u], from Lifetime


BeamLoss Detectors [a.u]


 $\times 10^6$ Halo Intensity [a.u]


normalized curves of Halo, Loss-Rate (from Lifetime), BeamLossDetectors and Vertical Beamsize

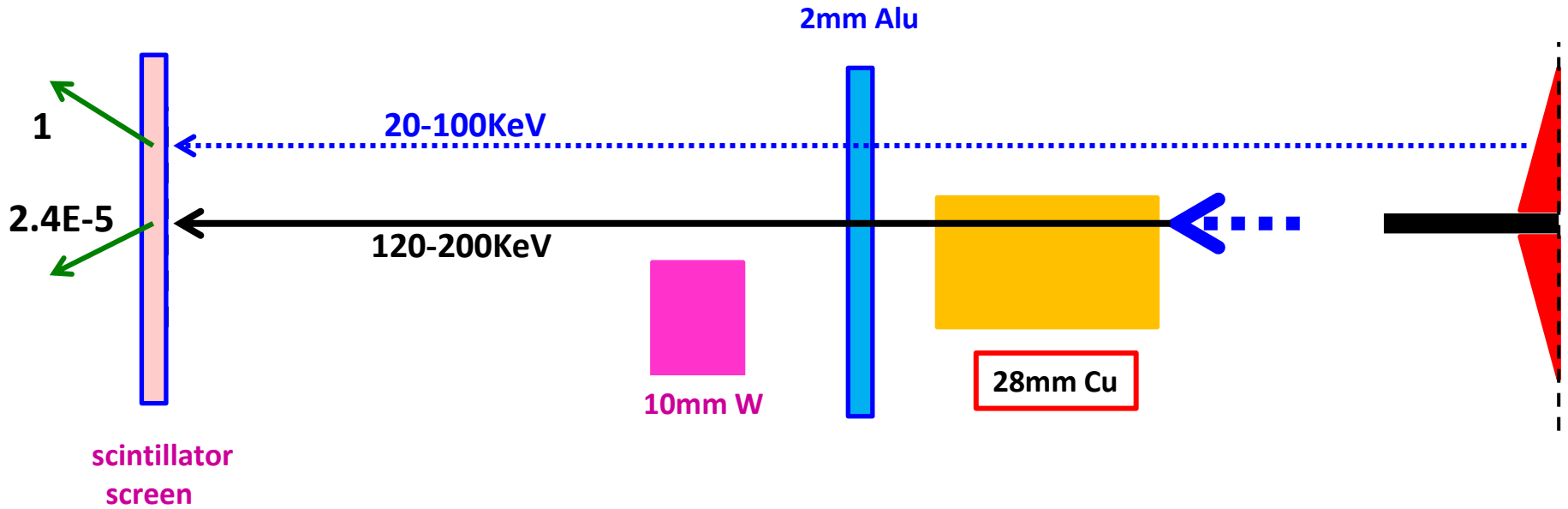
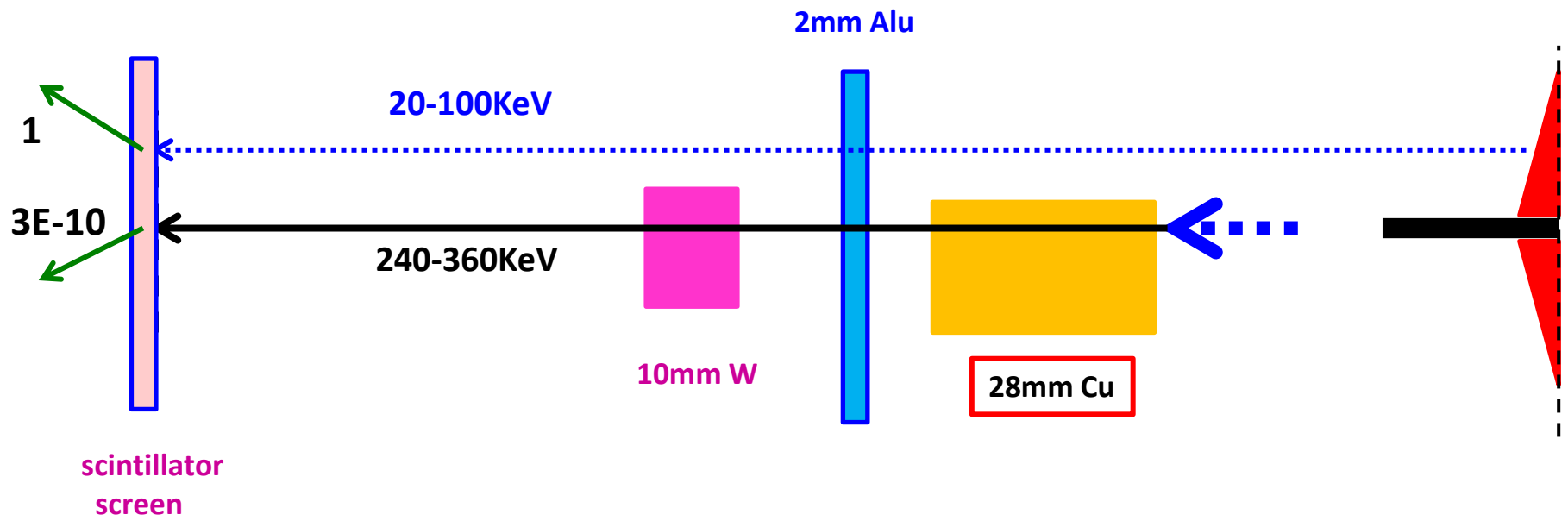


increasing beam-size, away from Touchek dominated scattering

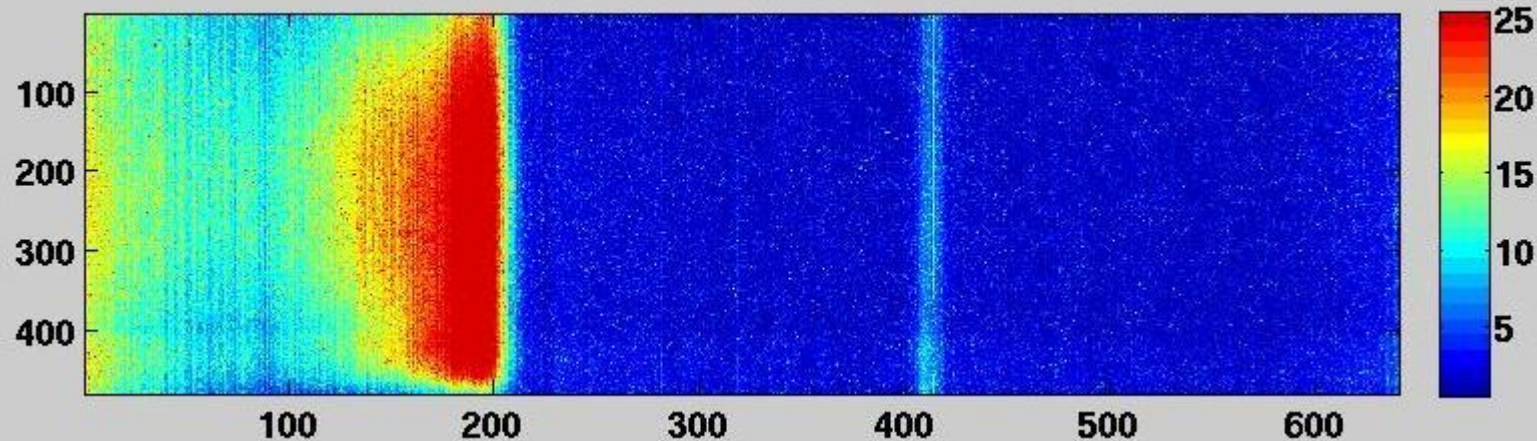
1) absolute calibration

2) limitations of spatial resolution & range of detection

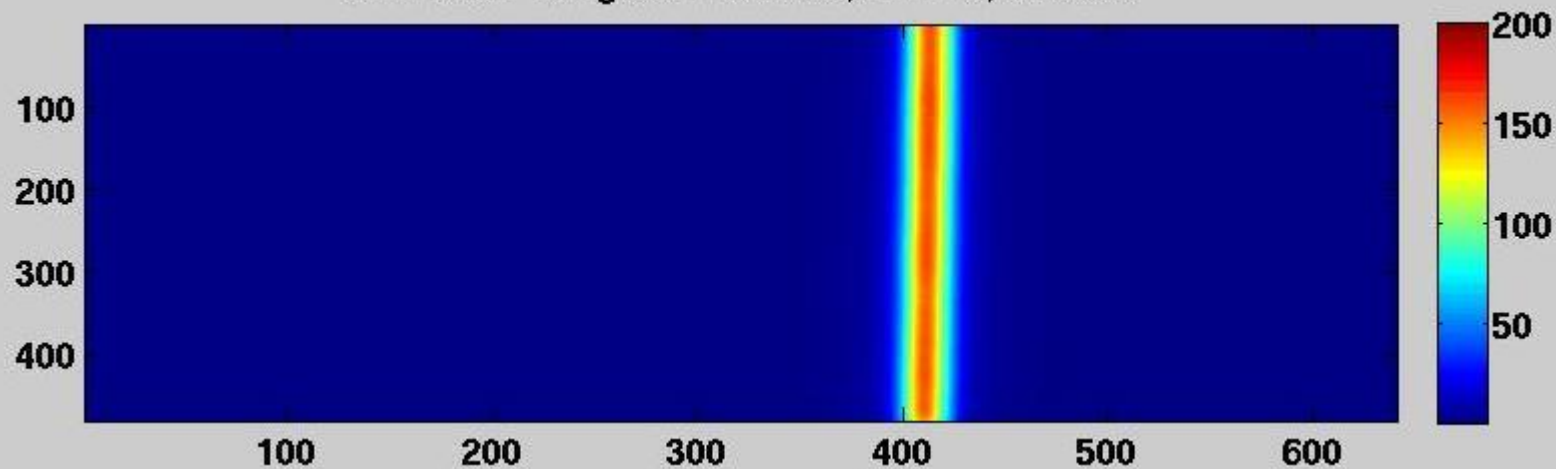
3) some initial problems resolved



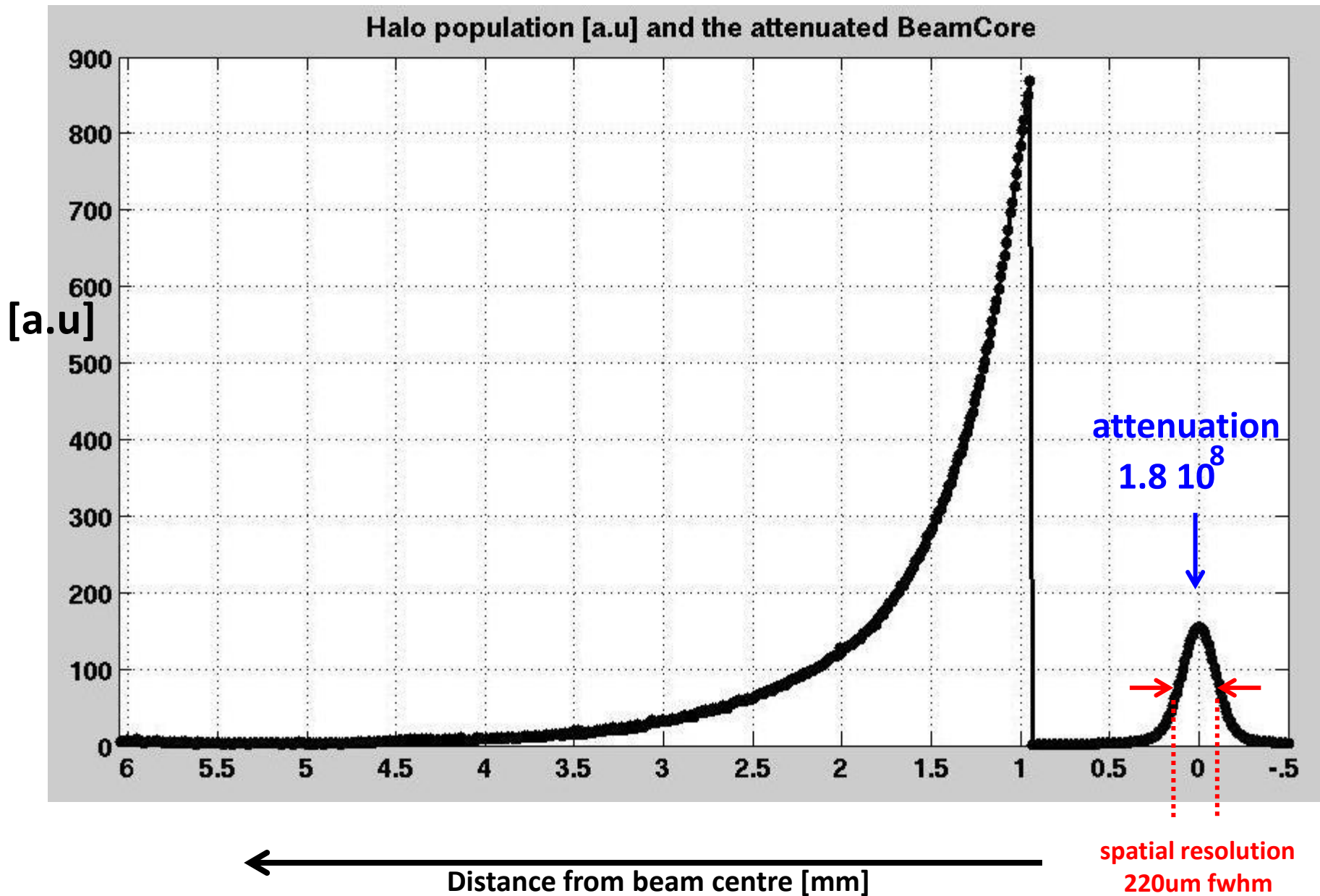
WITH 10mm Tungsten Blocker, T=130ms, G=30dB



WITHOUT Tungsten Blocker, T=1ms, G=0dB



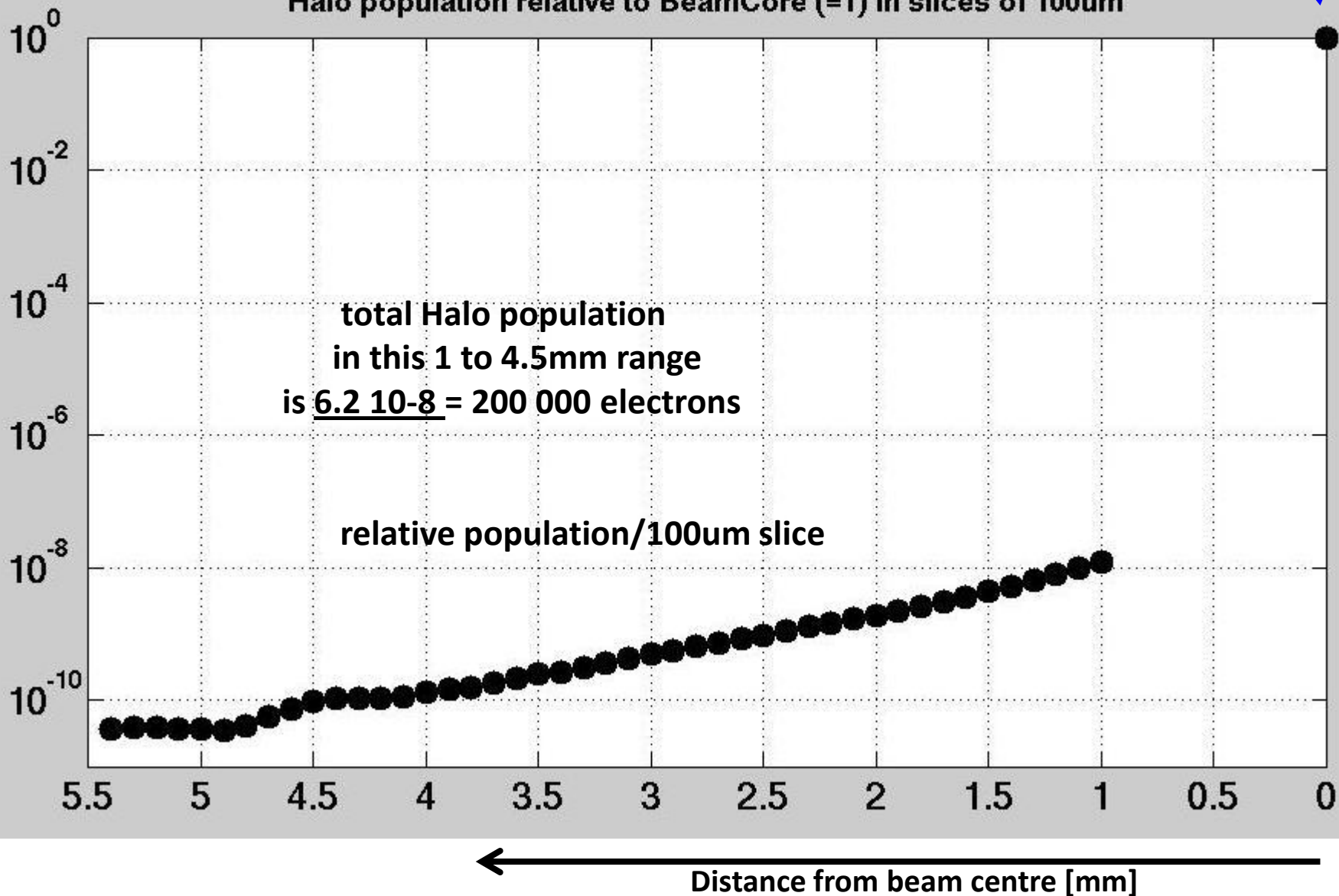
Attenuation = 28mm Cu & 30dB camera gain & 130 exposure time



BeamCore = 1



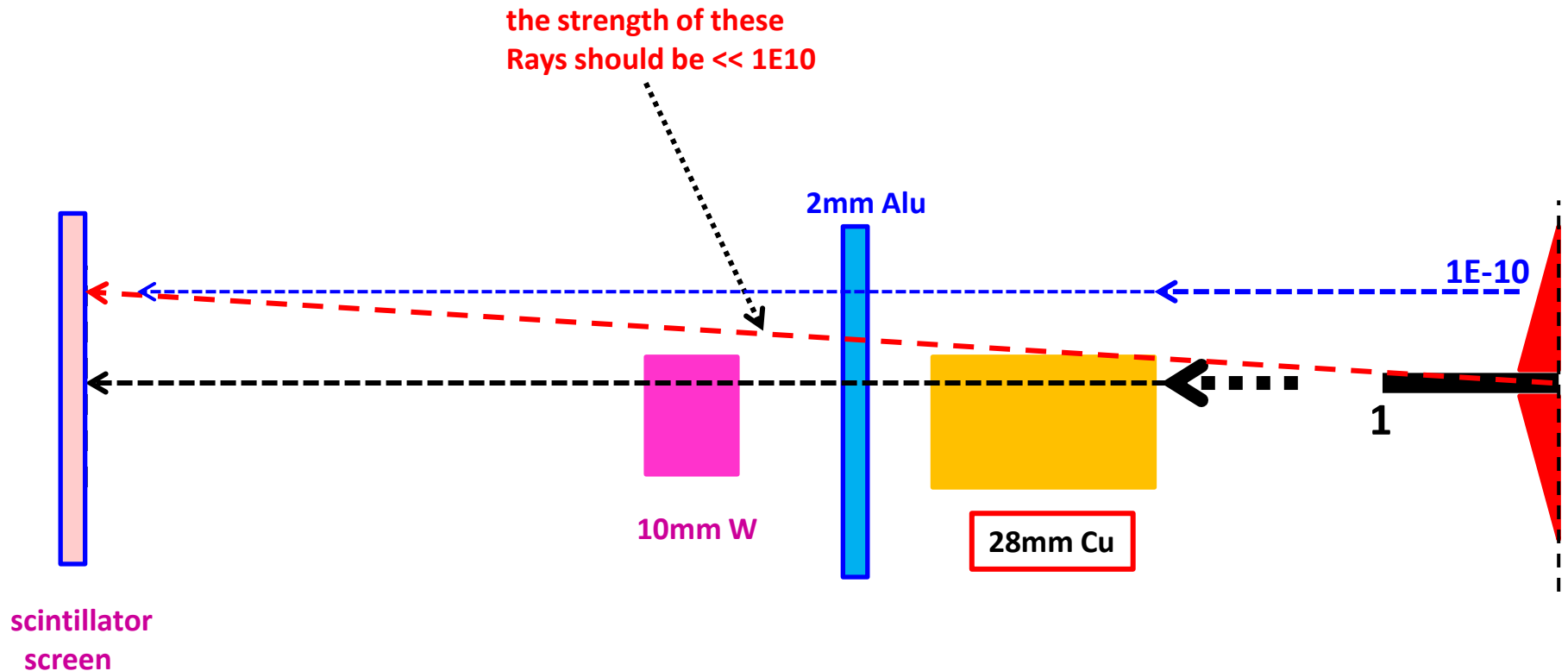
Halo population relative to BeamCore (=1) in slices of 100um



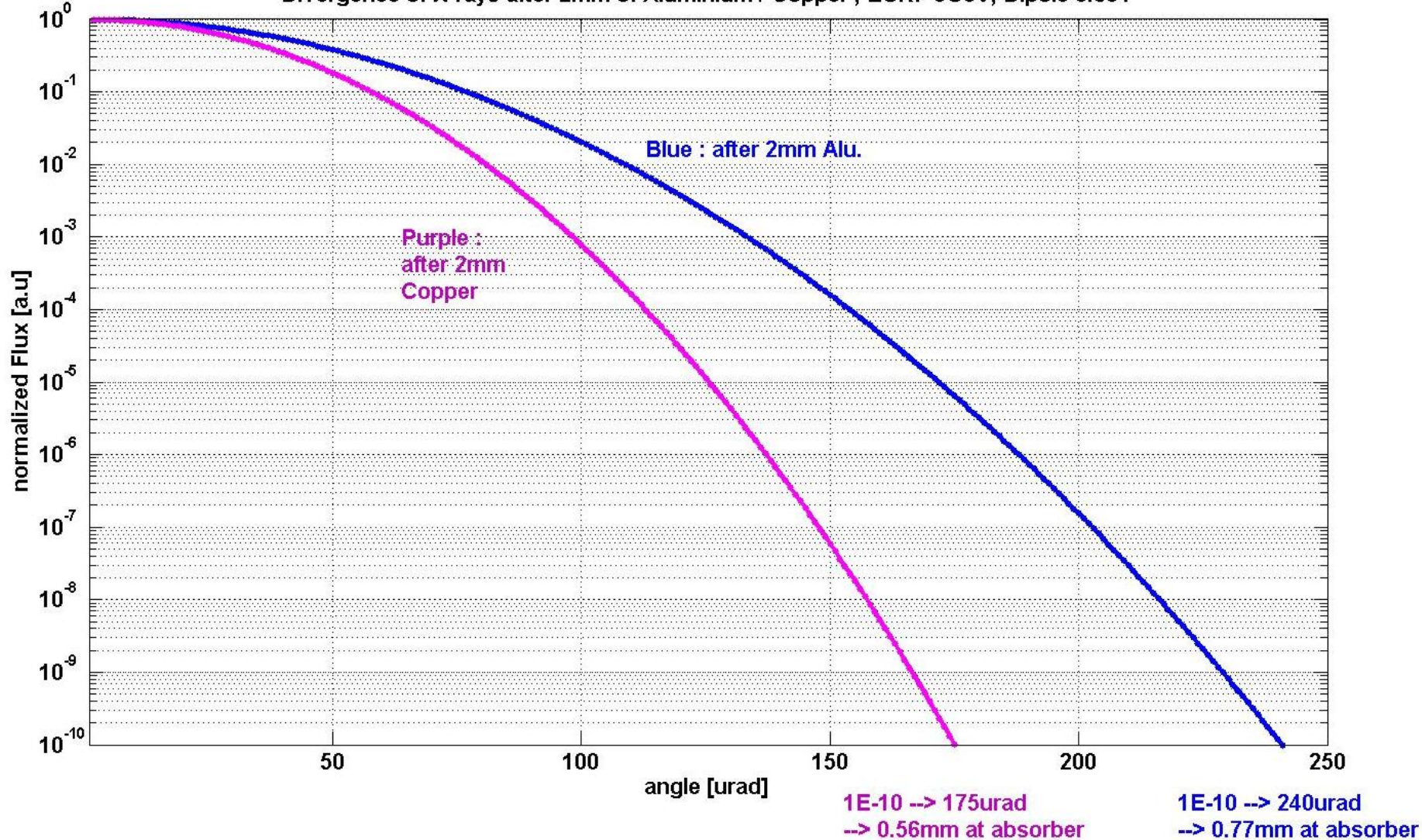
the “not-so-small” divergence of the “not-so-hard” X-rays

determining :

- 1) the spatial resolution of the monitor
- 2) how close to the beam-core we can measure

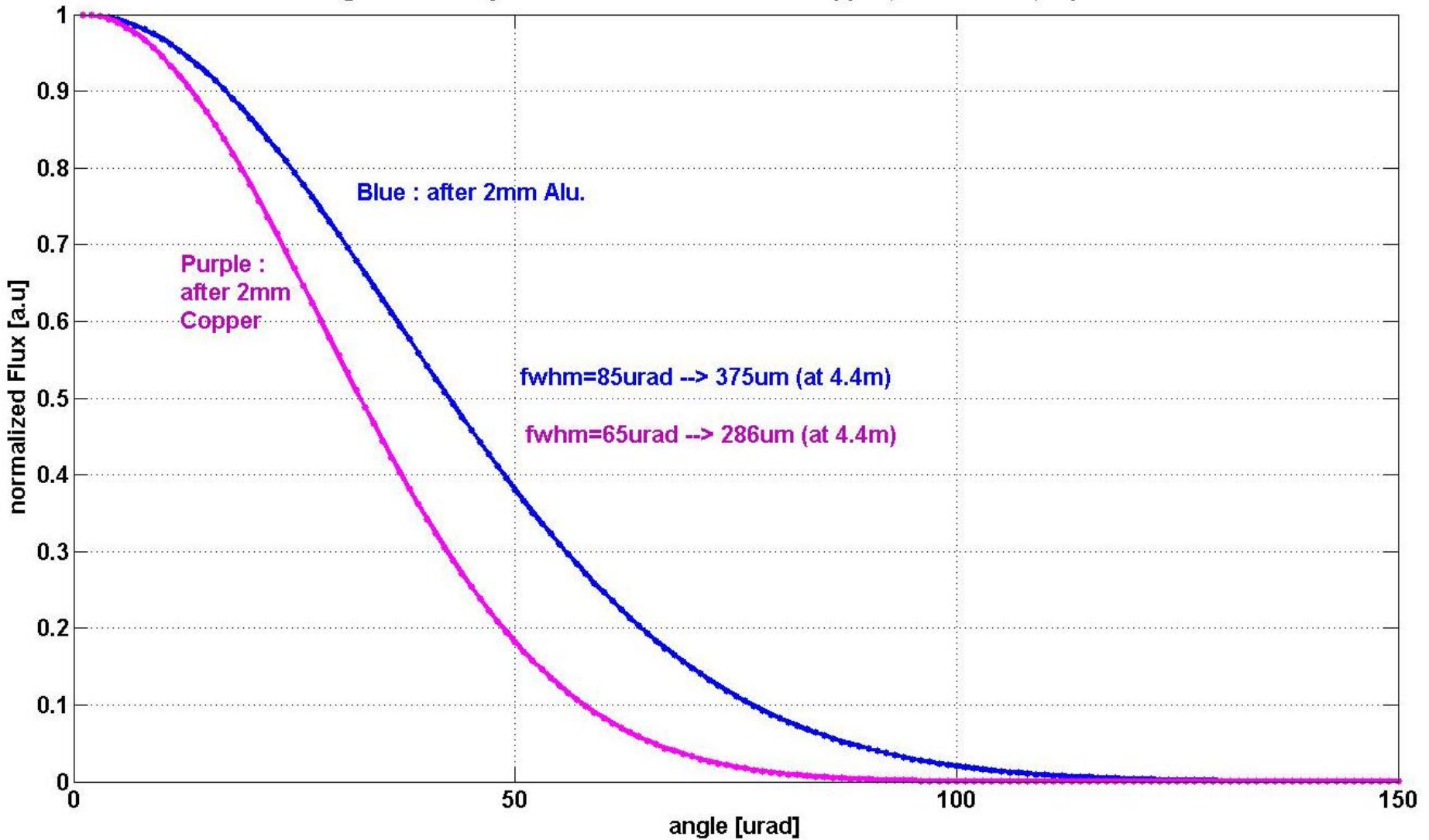


Divergence of X-rays after 2mm of Aluminium / Copper , ESRF 6GeV, Dipole 0.86T



but nearly a factor 6 less flux
with 2mm Copper

Divergence of X-rays after 2mm of Aluminium / Copper , ESRF 6GeV, Dipole 0.86T



but nearly a factor 6 less flux
with 2mm Copper

Some technical problems & challenges being solved :

phantom signals :

from scattering / reflections of the intense X-ray beamcore
from some of the (unavoidable) absorbers

solution (empirical) : move absorber & detector to a different part ...

in today's optimized set-up the phantoms have disappeared !!

fast degradation of the detector (CCD camera)

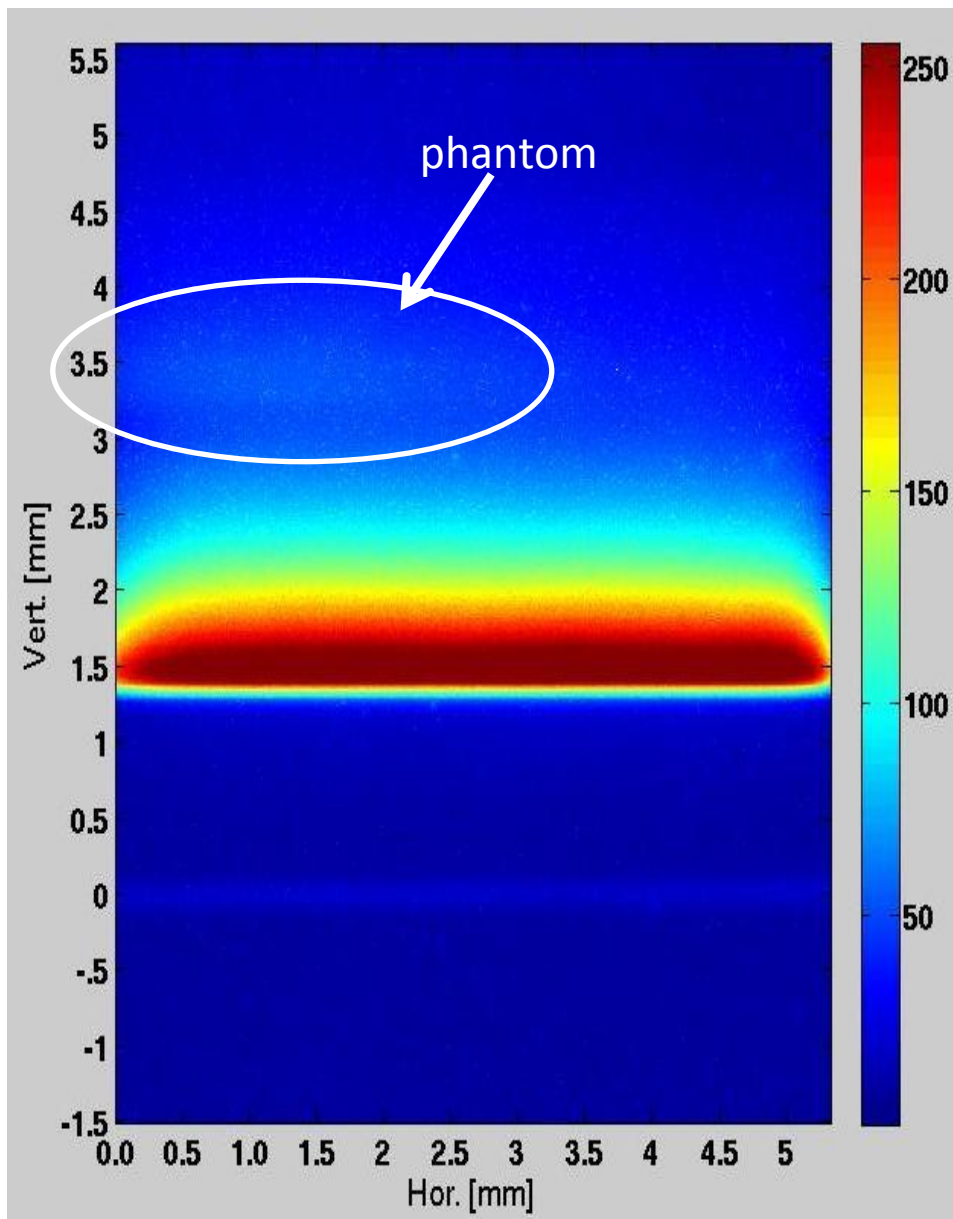
due to nasty radiation environment

used till recently : Flea camera (fire-wire standard) , 430 Euros

now : BlackFLY camera (GigE standard, power_over_Ethernet) , 240 Euros

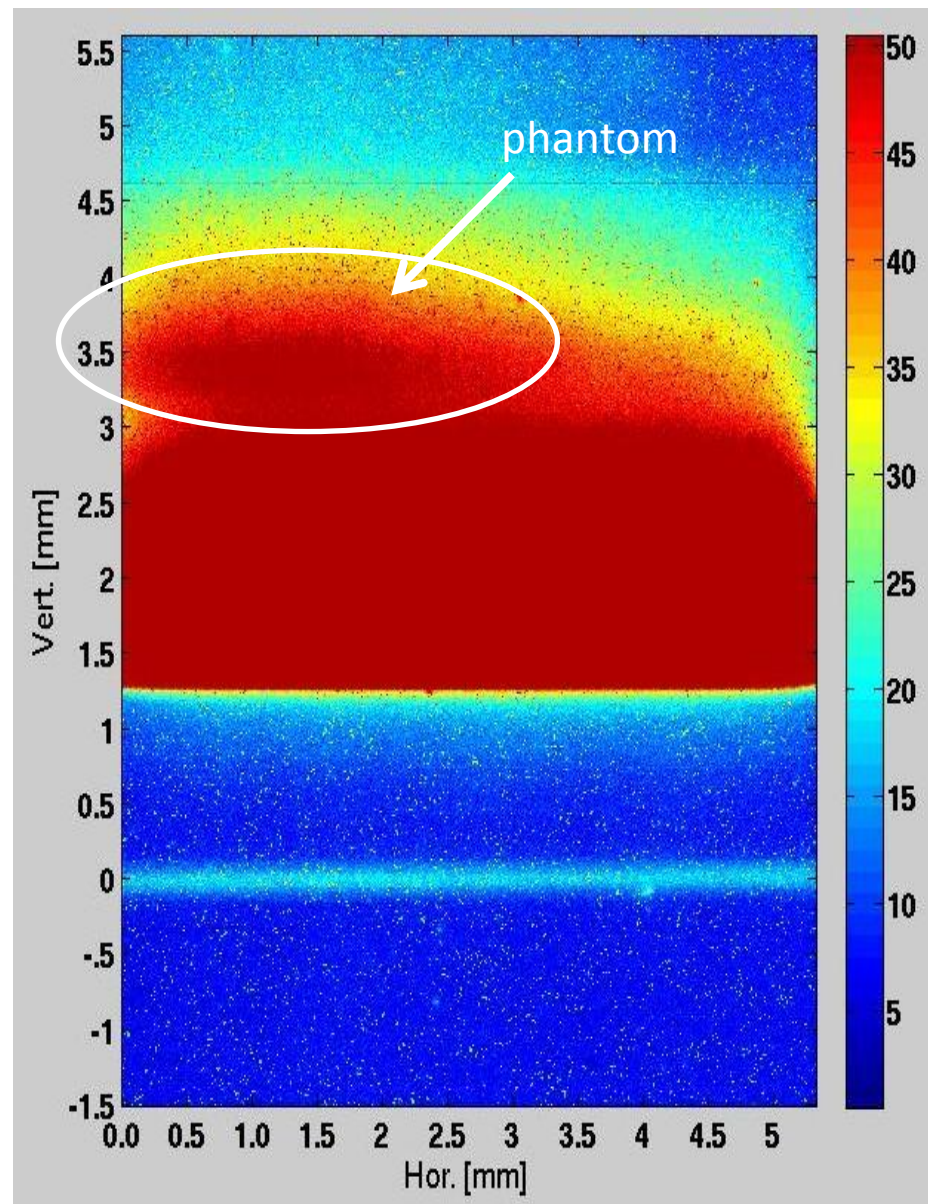
hardly any degradation noticeable after 2 weeks !!

Colormap
250



Hor. Accepted Dipole Radiation $\sim 1\text{mrad}$

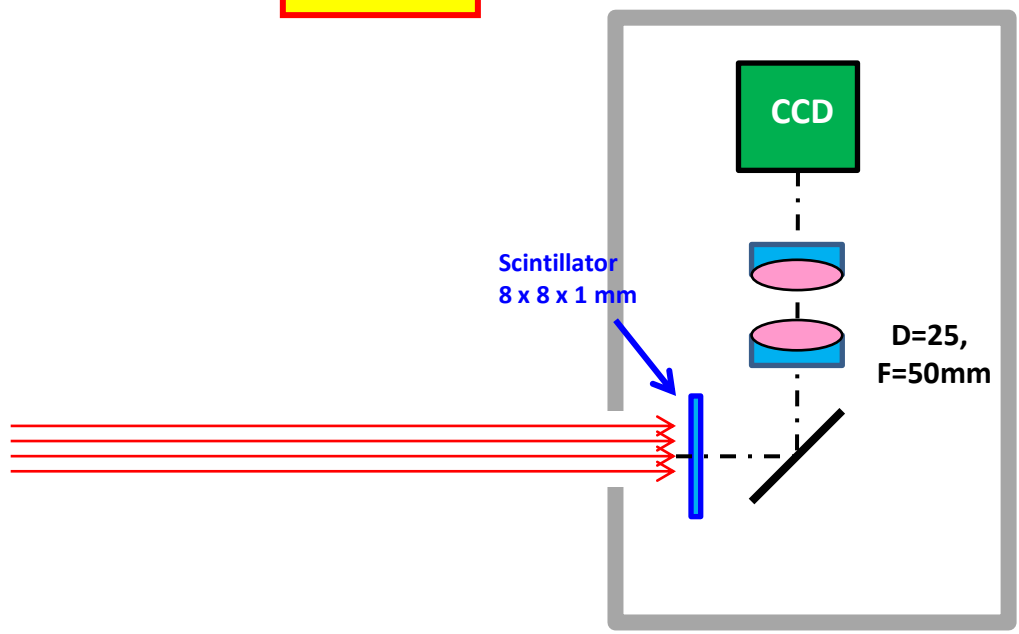
Colormap
50



Hor. Accepted Dipole Radiation $\sim 1\text{mrad}$

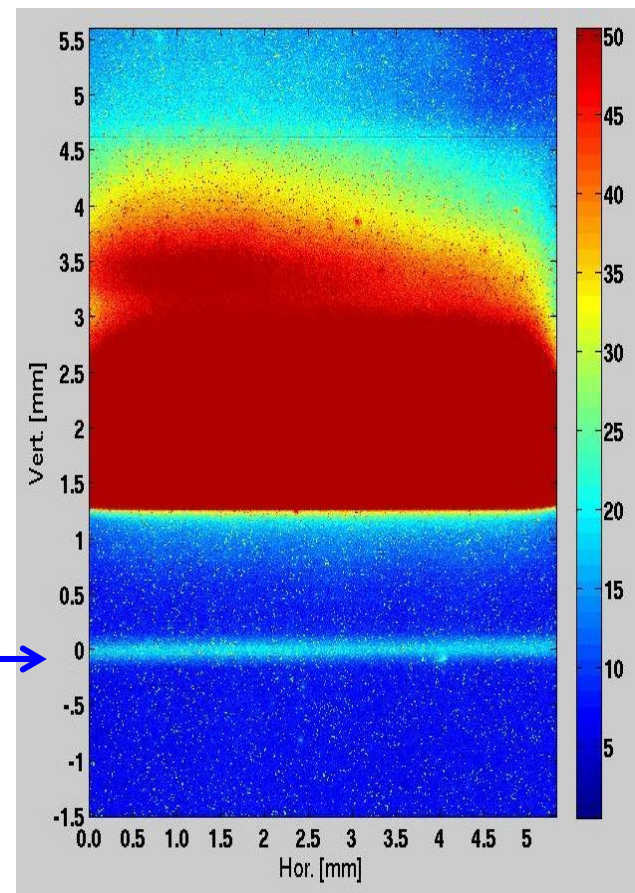
X-ray Imager

top view



No image intensifier, No cooled camera, just a cheap camera, at max gain (30dB) and exposure time (130millisec)

“degradation” means :
Loss of sensitivity & linearity
(mainly at high Gain)



Flea (430 Euros) 1/3" ICX424 Sony 7.4um
“degradation” within days ☹️

BlackFly (240 Euros) 1/3" ICX693 Sony 6um
much more resistant . . . 😊

Halo Iujah

